

Numerical Relativity - PHY 6938

Solutions to HW 5

1. Recall that two covariant derivative operators differ by a tensor C_{ik}^l , e.g.

$$D_i w_k = \bar{D}_i w_k - C_{ik}^l w_l.$$

Read Wald's book if you do not believe this.

Let's start with

$$R_{ijk}{}^l w_l = D_i D_j w_k - D_j D_i w_k$$

and note $D_i D_j w_k = \bar{D}_i (\bar{D}_j w_k - C_{jk}^m w_m) - C_{ij}^l (\bar{D}_l w_k - C_{lk}^m w_m) - C_{ik}^l (\bar{D}_j w_l - C_{jl}^m w_m)$.

Thus $R_{ijk}{}^l w_l = D_i D_j w_k - D_j D_i w_k = \bar{R}_{ijk}{}^l w_l + (2\bar{D}_{[j} C_{i]k}^l + 2C_{k[i}^m C_{j]m}^l) w_l$.

If $D_i \gamma_{kl} = 0$ and $\bar{D}_i \bar{\gamma}_{kl} = 0$, we have

$$C_{ij}^k = \frac{1}{2} \gamma^{kl} (\bar{D}_j \gamma_{li} + \bar{D}_i \gamma_{lj} - \bar{D}_l \gamma_{ij}).$$

We use this for the case of $\gamma_{ij} = \psi^4 \bar{\gamma}_{ij}$, $\gamma^{ij} = \psi^{-4} \bar{\gamma}^{ij}$, and find $\bar{D}_i \gamma_{jk} = 4\psi^3 \bar{\gamma}_{jk} \bar{D}_i \psi = 4\gamma_{jk} D_i \ln \psi$, which leads to

$$C_{ij}^l = (4\delta_{[i}^l \bar{D}_{j]} - 2\bar{\gamma}_{ij} \bar{\gamma}^{kl} \bar{D}_k) \ln \psi.$$

Then the Riemann tensor is $R_{ijk}{}^l = \bar{R}_{ijk}{}^l + 2\bar{D}_{[j} C_{i]k}^l + 2C_{k[i}^m C_{j]m}^l = \bar{R}_{ijk}{}^l + 4\delta_{[i}^l \bar{D}_{j]} \bar{D}_k \ln \psi - 4\bar{\gamma}^{lm} \bar{\gamma}_{k[i} \bar{D}_{j]} \bar{D}_m \ln \psi + 8(\bar{D}_{[i} \ln \psi) \delta_{j]}^l \bar{D}_k \ln \psi - 8(\bar{D}_{[i} \ln \psi) \bar{\gamma}_{j]k} \bar{\gamma}^{lm} \bar{D}_m \ln \psi - 8\bar{\gamma}_{k[i} \delta_{j]}^l \bar{\gamma}^{mn} (\bar{D}_m \ln \psi) \bar{D}_n \ln \psi$. So the Ricci tensor becomes $R_{ik} = R_{ilk}{}^l = \bar{R}_{ik} - 2\bar{D}_i \bar{D}_k \ln \psi - 2\bar{\gamma}_{ik} \bar{\gamma}^{lm} \bar{D}_l \bar{D}_m \ln \psi + 4(\bar{D}_i \ln \psi) \bar{D}_k \ln \psi - 4\bar{\gamma}_{ik} \bar{\gamma}^{lm} (\bar{D}_l \ln \psi) \bar{D}_m \ln \psi$.

Form this we can compute the Ricci scalar and find

$$R = \psi^{-4} \bar{R} - 8\psi^{-5} \bar{D}_k \bar{D}^k \psi.$$

[Note: $\bar{D}_i \bar{D}_j \ln \psi = \bar{D}_i (\psi^{-1} \bar{D}_j \psi) = -\psi^{-2} (\bar{D}_i \psi) \bar{D}_j \psi + \psi^{-1} \bar{D}_i \bar{D}_j \psi$. So $\psi^{-1} \bar{D}_i \bar{D}_j \psi = \bar{D}_i \bar{D}_j \ln \psi + \psi^{-2} (\bar{D}_i \psi) \bar{D}_j \psi = \bar{D}_i \bar{D}_j \ln \psi + (\bar{D}_i \ln \psi) (\bar{D}_j \ln \psi)$.]

2. Let's start with $D_j (\psi^n S^{ij}) = n\psi^{n-1} S^{ij} D_j \psi + \psi^n D_j S^{ij}$ and replace D_j by $\bar{D}_j$ and C_{ij}^l as in 1.

Take into account that S^{ij} is symmetric. We then find

$$D_j (\psi^n S^{ij}) = \psi^n [\bar{D}_j S^{ij} + (10 + n) S^{ij} \bar{D}_j \ln \psi - 2\bar{\gamma}_{jk} S^{jk} \bar{D}^i \ln \psi].$$

3. Since $(LW)^{ij} := D^i W^j + D^j W^i - (2/3) \gamma^{ij} D_k W^k$, we need $D^i W^j = \gamma^{im} D_m W^j = \gamma^{im} (\bar{D}_m W^j + C_{mk}^j W^k)$.

With the C_{mk}^j from 1. we find:

$$(LW)^{ij} = \psi^{-4} (\bar{L}W)^{ij}$$

4. We have $\gamma_{ij} = \psi^4 \delta_{ij}$. Thus $\bar{D}_i = \partial_i$. Now from 3. we find $(L\beta)^{ij} = \psi^{-4}(\bar{L}\beta)^{ij} = \psi^{-4}(\partial^i \beta^j + \partial^j \beta^i - (2/3)\delta^{ij} \partial_k \beta^k)$.

Here we have $\beta^i = \epsilon^{ijk} \Omega^j x^k$ (where we sum over j and k). Thus $(L\beta)^{ij} = 0$.

5. If $\gamma_{ij} = \psi^4 \delta_{ij}$ the ADM mass is given by

$$M_{ADM} = -\frac{1}{2\pi} \lim_{r \rightarrow \infty} \oint_S \partial_j \psi dS^j.$$

If we construct puncture initial data and solve the Hamiltonian constraint for u we obtain $\psi = 1 + \sum_A \frac{m_A}{2r_A} + u$ as conformal factor. This results in

$$M_{ADM} = \sum_A m_A - \frac{1}{2\pi} \lim_{r \rightarrow \infty} \oint_S \partial_j u dS^j = \sum_A m_A - \frac{1}{2\pi} \lim_{r \rightarrow \infty} \int_0^\pi \int_0^{2\pi} (\partial_r u) \sin \theta \, d\theta d\phi.$$

[Im letzten Integral fehlt ein Faktor r^2 .]